

# The Neutrino Landscape

We have come a long way since the Solar Neutrino Problem  
still much to do

Bright Future: see “Neutrinos” White Paper

Conveners: A. de Gouvea, K. Pitts, K. Scholberg, G.P. Zeller

**William J. Marciano**

**WINP**

**Workshop on the Intermediate Neutrino Program**

**Feb 4, 2015**



# **Neutrino Masses & Mixing: The Heroic Years**

**R. Davis** (BNL Chemist) – The Solar Neutrino Problem  
~1/3 expected solar  $\nu_e$  flux  
>30 yr. **Homestake** Effort

Kamiokande & IMB Collaborations – **Supernova 1987A** &  
Hints of Atm.  $\nu_\mu$  osc. (depletion)

**M. Koshiba** (Super Kamiokande) finds  $\nu_\mu$ - $\nu_\tau$  atm osc. Large!

**SuperK, SNO, Kamland**, Gallium Exps,..K2K, MINOS, T2K  
Daya Bay, RENO, Double Chooz... **Neutrino Osc. Matures**  
**Large 3 Generation Mixing Paradigm Cofirmed!**

|  $\Delta m_{32}^2$  | &  $\Delta m_{21}^2$  determined!

### 3 Generation Mixing Formalism

$$\begin{pmatrix} |\nu_e\rangle \\ |\nu_\mu\rangle \\ |\nu_\tau\rangle \end{pmatrix} = U \begin{pmatrix} |\nu_1\rangle \\ |\nu_2\rangle \\ |\nu_3\rangle \end{pmatrix} \quad (1)$$

$$U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix}$$

$$c_{ij} = \cos \theta_{ij} \quad , \quad s_{ij} = \sin \theta_{ij}$$

$$J_{CP} \equiv \frac{1}{8} \sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23} \cos \theta_{13} \sin \delta. \quad (2)$$

# Current Neutrino Mass & Mixing Parameters

- $\Delta m_{32}^2 = m_3^2 - m_2^2 = \pm 2.4(1) \times 10^{-3} \text{ eV}^2$   
(atmospheric, beam & reactor all precise & consistent!)
- $\Delta m_{21}^2 = m_2^2 - m_1^2 = +7.5(2) \times 10^{-5} \text{ eV}^2$  (Solar, Kamland)

**$|\Delta m_{21}^2 / \Delta m_{32}^2 \approx 1/30| \rightarrow \text{CP Violation Exp Doable!}$**

Hierarchy  $m_3 > m_1 \& m_2$  (normal) or  $m_3 < m_1 \& m_2$  (inverted)?

## **Large Mixing!**

$$\theta_{23} \sim 38 \pm 1^\circ \quad \sin 2\theta_{23} \approx 0.97(1) \quad (\theta_{23} \text{ or } 90^\circ - \theta_{23}) \quad (\text{atm.})$$

$$\theta_{12} \sim 34 \pm 1^\circ \quad \sin 2\theta_{12} = 0.93(2) \quad (\text{solar})$$

$$\theta_{13} \leq 9.0 \pm 0.5^\circ \quad \sin 2\theta_{13} = 0.31(1) \quad (\text{reactor})$$

$$0 \leq \delta \leq 360^\circ ? \quad (\text{some recent sensitivity})$$

$$J_{\text{CP}} \sim 0.03 \sin \delta \quad (\text{potentially large!}) \quad \text{CKM} \sim 2 \times 10^{-5}$$

### 3 Generation $\nu_\mu$ - $\nu_e$ Oscillations

$$P(\nu_\mu \rightarrow \nu_e) = P_I(\nu_\mu \rightarrow \nu_e) + P_{II}(\nu_\mu \rightarrow \nu_e) + P_{III}(\nu_\mu \rightarrow \nu_e) \\ + \text{matter} + \text{smaller terms}$$

$$\mathbf{P}_I(\nu_\mu \rightarrow \nu_e) = \sin^2 \theta_{23} \sin^2 2\theta_{13} \sin^2 \left( \frac{\Delta m_{31}^2 L}{4E_\nu} \right)$$

$$\mathbf{P}_{II}(\nu_\mu \rightarrow \nu_e) = \frac{1}{2} \sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23} \cos \theta_{13} \\ \sin \left( \frac{\Delta m_{21}^2 L}{2E_\nu} \right) \times \left[ \sin \delta \sin^2 \left( \frac{\Delta m_{31}^2 L}{4E_\nu} \right) \right. \\ \left. + \cos \delta \sin \left( \frac{\Delta m_{31}^2 L}{4E_\nu} \right) \cos \left( \frac{\Delta m_{31}^2 L}{4E_\nu} \right) \right]$$

$$\mathbf{P}_{III}(\nu_\mu \rightarrow \nu_e) = \sin^2 2\theta_{12} \cos^2 \theta_{13} \cos^2 \theta_{23} \sin^2 \left( \frac{\Delta m_{21}^2 L}{4E_\nu} \right)$$

For antineutrinos,  $\delta \rightarrow -\delta$  and opposite matter effect.

# CP Violation Asymmetry

$$A_{CP} \equiv \frac{P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)}{P(\nu_\mu \rightarrow \nu_e) + P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)} \quad (3)$$

To leading order in  $\Delta m_{21}^2$  ( $\sin^2 2\theta_{13}$  is not too small):

$$A_{CP} \simeq \frac{\cos \theta_{23} \sin 2\theta_{12} \sin \delta}{\sin \theta_{23} \sin \theta_{13}} \left( \frac{\Delta m_{21}^2 L}{4E_\nu} \right) + \text{matter effects} \quad (4)$$

$$F.O.M. = \left( \frac{\delta A_{CP}}{A_{CP}} \right)^{-2} = \frac{A_{CP}^2 N}{1 - A_{CP}^2} \quad (5)$$

$N$  is the total number of  $\nu_\mu \rightarrow \nu_e + \bar{\nu}_\mu \rightarrow \bar{\nu}_e$  events. Since  $N$  falls (roughly) as  $\sin^2 \theta_{13}$  and  $A_{CP}^2 \sim 1/\sin^2 \theta_{13}$ , to a first approximation the F.O.M. is independent of  $\sin \theta_{13}$ . Similarly, given  $E_\nu$  the neutrino flux and consequently  $N$  falls as  $1/L^2$  but that is canceled by  $L^2$  in  $A_{CP}^2$ .

## CP Violation Insensitivities

- To a very good approx., our statistical ability to determine  $\delta$  or  $A_{cp}$  is independent of  $\sin^2 2\theta_{13}$  (down to  $\sim 0.003$ ) and the detector distance  $L$  (for long distance).

It turned out  $\sin^2 2\theta_{13} \approx 0.1!$

about 2-3 times larger than assumed in early studies

***precision  $\theta_{13}$  &  $\delta$  determination easier!***

Smaller  $A_{CP}$  Might help some systematics

## *What do we still need to learn?*

1. **Sgn  $\Delta m_{32}^2$ ?** (Important for Neutrinoless  $\beta\beta$  Decay)  
Earth Matter Effect (SM or “New Physics”)  
or Reactor Precision (Possible Near Term)
- 2. **Value of  $\delta$ ?,  $J_{CP}$ ?, CP Violation?** *(Holy Grail)*  
*Relationship to Leptogenesis*
- 3. **Precision  $\Delta m_{32}^2$ ,  $\Delta m_{21}^2$ ,  $\theta_{23}$ ,  $\theta_{12}$ ,  $\theta_{13}$**  (better than 1%!)  
Redundancy & neutrinos vs antineutrinos  
Unitarity Violation? – Sterile neutrino Mixing
- 4. **“New Physics”** - Sterile  $\nu$ , **Very Weak** Long/Short  
Distance New Physics (*The Dark World?*)...



## **CP Violation Requirements & LBNF**

- What does it take to measure  $\delta$  to  $\pm 15^\circ$  in about 6 yrs.?

Answer (Approx.): 200kton Water Cerenkov Detector

Approx 20% Acceptance or

40 kton LArgon 90% Acceptance

or Hybrid combination

+ Traditional Horn Focused  $\nu_\mu$  WBB (0.5-5GeV)

powered by

1-2MW proton accelerator

Long Distance > 1000km

Other approaches eg. Neutrino factory, low energy beams...

## **Exp. Goals: LBNF & Near Term Efforts**

- **Measure Leptonic CP Violation & phase  $\delta$**   
**Our Origin (Leptogenesis Matter-antiMatter Asy.)**

**Determine (Precisely) Neutrino Mass & Mixing Parameters**

**Redundant Comparison Tests (Unitarity)**

**Understand Pattern (Underlying Symmetry eg A4)**

**Search for “New Physics” eg very weakly coupled new long or short distance effects (in Matter) vs MSW**

**Search for Sterile Neutrinos (small mixing)**

## ***Other Physics***

- ***Atm. vs Beam Neutrino Oscillations***
- ***Supernova Neutrinos (Relic & New),***
- ***Proton Decay ( $B-L=0$ )... Neutron-antiNeutron Osc.  $\Delta B=2$***

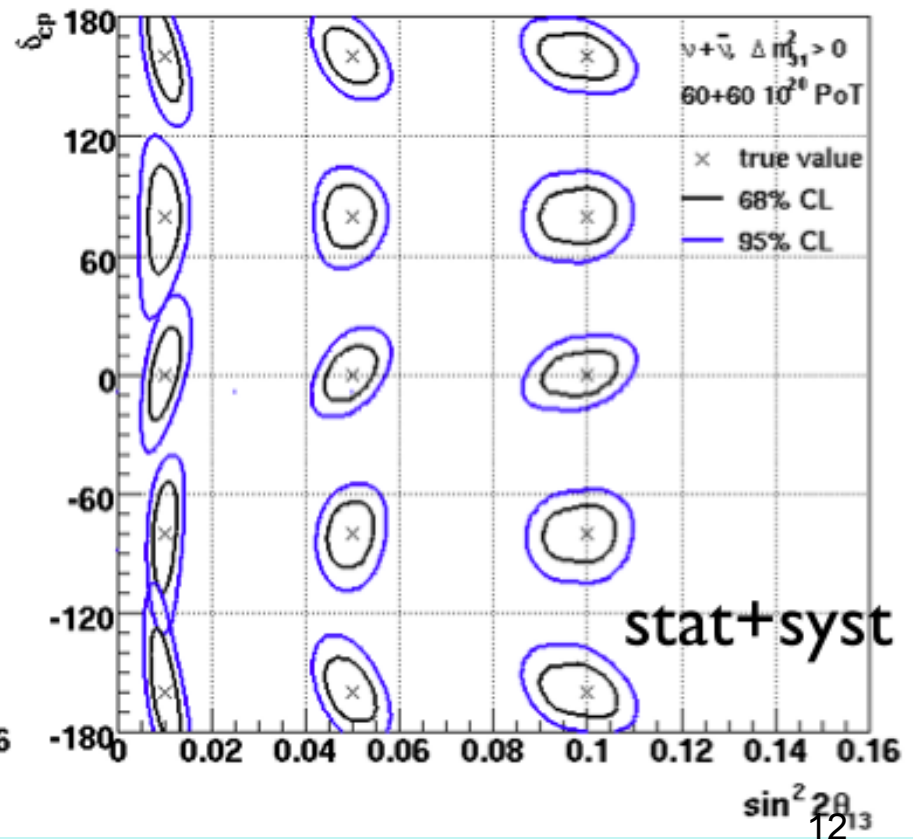
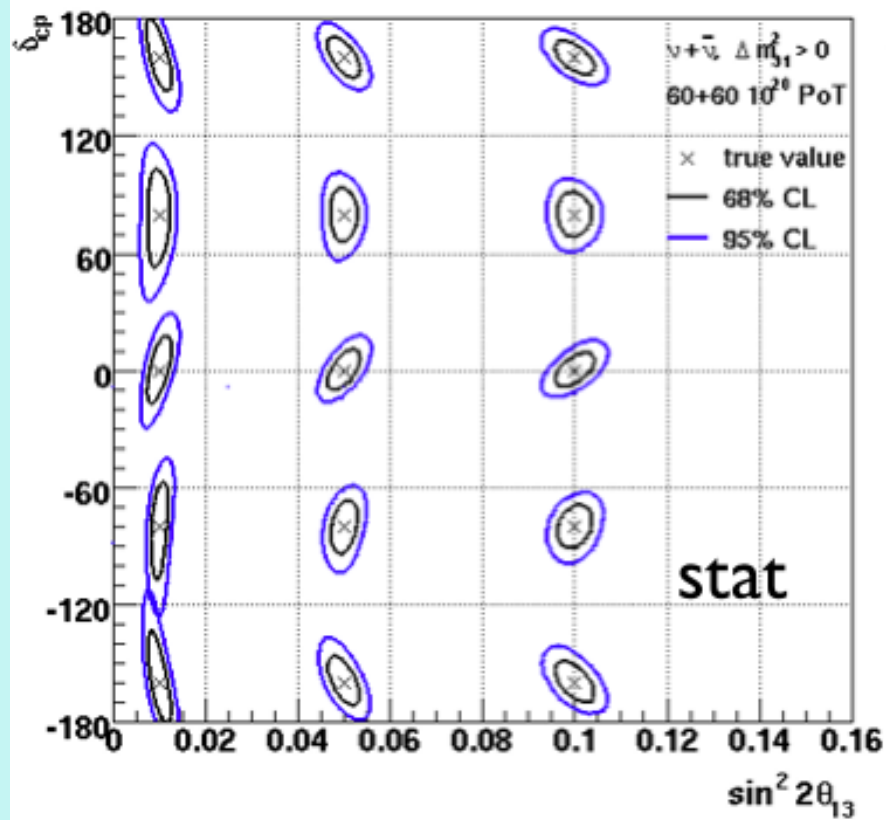
## ***Broad Revolutionary Discovery Potential***

***Big Underground Detectors Originally Proposed for  
Proton Decay Searches Motivated by Grand Unification  
Facility for Revolutionary Discoveries***

# CP Phase & $\theta_{13}$ LBNF Sensitivity

WCC 1300 km 300kT

(-95% CL -68% CL)



$\Delta m^2_{32}$  and  $\sin^2 2\theta_{32}$  can be measured in long baselines as functions of  $E_\nu$  (also obtained from atmospheric  $\nu$ ).

**$\nu_\mu \rightarrow \nu_\mu$  &  $\text{anti}\nu_\mu \rightarrow \text{anti}\nu_\mu$  Comparison**

Usually phrased as a test of CPT (true in vacuum)

Apparent CPT violation  $\rightarrow$  “New Physics” in  $\nu$  interactions  
(in matter)

$$\varepsilon \sqrt{2} G_F \nu \gamma_\mu \nu' f \gamma^\mu f, \quad f=e, u, d$$

Potential changes sign  $\nu_\mu \rightarrow \text{anti}\nu_\mu$

Sterile Neutrinos? etc

## Goals

**Long Term: CP violation in neutrino oscillations**

***LBNE: 1300km, WBB, 1-2MW, 40kton LAr, 10yrs***

Proton Decay ( $10^{35}\text{yr}$ )

Similar Detector Requirements (Fortuitous)

***Also: Atm & supernova  $\nu$ , neutron-antineutron osc.,...***

### **Nearer Term & LBNF**

**Sgn  $\Delta m_{32}^2$ ?** (Important for Neutrinoless  $\beta\beta$  Decay)

**Precision  $\Delta m_{32}^2$ ,  $\Delta m_{21}^2$ ,  $\theta_{23}$ ,  $\theta_{12}$ ,  $\theta_{13}$**  (goal?  $\pm 1\%$ !)

**“New Physics” - Sterile  $\nu$ , Very Weak New Interactions...Neutrino-antineutrino differences?**

***Anticipate (Hope For) Surprises***

## **Lessons Learned for Future Neutrino Experiments**

- Understand all backgrounds before starting  
(Recent experiments very good)
- Over design detector and its capabilities if possible  
(little room for descoping)

Definitive goals: eg. Sterile neutrino sensitivity

Start as soon as possible (long time commitment)